

CALLISTO status report/newsletter #100

This report is number 100 after the 1st one on July 4th 2003 thus, this project is now running for 21 years and the network is still growing.

Several recent 1st lights



Fig. 1: 1st light from new station Strassolt in Italy Contact: Alex Marassi <alessandro.marassi(at)inaf.it>



Fig. 2: 1st light from new station Taiwan-NCU Contact: Ya-Hui Yang <yhyang(at)jupiter.ss.ncu.edu.tw>

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Fig. 4: 1st light from station ESSEN in Germany after years of maintenance break. Antenna biconical. Contact: Jochen Plessmann <do1jpa(at)gmx.de>





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Fig. 5: Group photo with LWA in Chiapas, Mexico. Contact: Ernesto Aguilar Rodríguez <Ernesto(at)igeofisica.unam.mx>





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Fig. 5: Logarithmic periodic dipole array (LPDA) including front-end box containing low noise amplifier. Antenna with tracking system. Station: Croatia-Visnjan Contact: Marko Radolović <mradolovic(at)gmail.com>



e-Callisto burst statistics October 2024



Fig. 6: Compilation of all visually detected bursts from all Callisto-stations which provide data to the e-Callisto network. There are clear 'winners' of the May-'competition', Australia-ASSA. Still eagerly looking for an AI-solution to automatically generate a burst-list and to save many hours day by day to perform this rather boring job, visual inspection of thousands of FIT-files.

Last 4 burst-plots are always available here: https://e-callisto.org/Data/data.html







CESRA NEWS

Data Release of Solar Radio Bursts observed by CBSm at the metric wavelength by Yao Chen et al. https://www.astro.gla.ac.uk/users/eduard/cesra/?p=3773

Energetics of compressive waves in the solar corona by Francesco Azzollini et al https://www.astro.gla.ac.uk/users/eduard/cesra/?p=3783

https://www.astro.gla.ac.uk/users/eduard/cesra/?p=3809

The angular and frequency dependence of solar radio burst rise and decay times using multi-spacecraft observations by Nicolina Chrysaphi et al.

https://www.astro.gla.ac.uk/users/eduard/cesra/?p=3852

Time-Profile Study of Type III Solar Radio Bursts Using Parker Solar Probe by Tulsi Thapa and Yihua Yan https://www.astro.gla.ac.uk/users/eduard/cesra/?p=3848

Periods and frequency drifts of groups of narrowband decimetric spikes by M. Karlický, J. Dudík and J. Rybák https://www.astro.gla.ac.uk/users/eduard/cesra/?p=3860

Generation of Series of Meter/Decimeter III type Bursts During Thermal Phase of Solar Flare by Meshalkina and Altyntsev https://www.astro.gla.ac.uk/users/eduard/cesra/?p=3884







Recent papers

Drone-Based Antenna Beam Calibration in the High Arctic https://arxiv.org/html/2407.00856v1

Regarding "Radio signature of the strong compression between a streamer and a coronal hole boundary" by Aguilar-Rodriguez et al DOI: 10.3847/2041-8213/ad631b The Astrophysical Journal Letters ad631b

Observation of an Extraordinary Type V Solar Radio Burst: Nonlinear Evolution of the Electron Two-Stream Instability Arnold O. Benz · Clemens R. Huber · Vincenzo Timmel · Christian Monstein https://doi.org/10.1007/s11207-024-02395-8

AOB

- If you have some stuff to present to the Callisto community, please let me know
- CALLISTO or Callisto denotes to the spectrometer itself while e-Callisto denotes to the worldwide network.



- General information and data access here: <u>https://e-callisto.org/</u>
- e-Callisto data are hosted at University of Applied Sciences, Institute for Data Science FHNW in Brugg/Windisch, Switzerland. Additionally, data are available at ESA site here: ESA Space Weather Portal (<u>https://swe.ssa.esa.int/</u>).
- University of Alcalá in Spain is also hosting e-Callisto data here: http://212.128.70.189/ From now on Bussons Gordo Javier <u>javier.bussons(at)uah.es</u> from Alcalá is the new Co-PI and will support my activities related to CALLISTO instrument and e-Callisto network.
- In case you (as the responsible person for operating and maintenance of Callisto) are leaving the institute or, if you are retiring, please send me name and email address of the successor.
- New product available, a calibration unit for antennas with direct access to the dipoles, such as CLP-5130. Calibration only for frequencies below 900 MHz (due to low cost components)





More information here (see bottom of this page): https://e-callisto.org/Products/Products.html



Fig. 7: Calibration unit containing semiconductor switch and a noise source. Unit will be supplied with a calibration table and free Python script to control the unit as well as to calibrate FIT-files from CALLISTO.

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Respond instead directly to me at: cmonstein(at)swissonline.ch or to the new Co-PI javier.bussons(at)uah.es

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